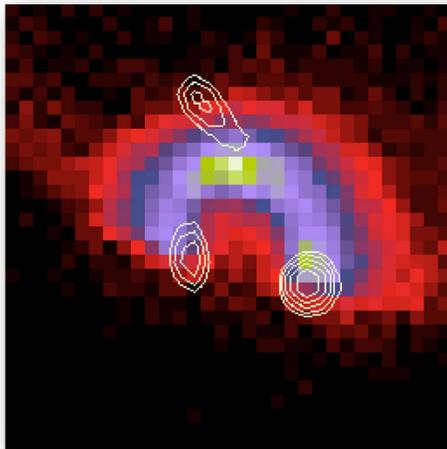
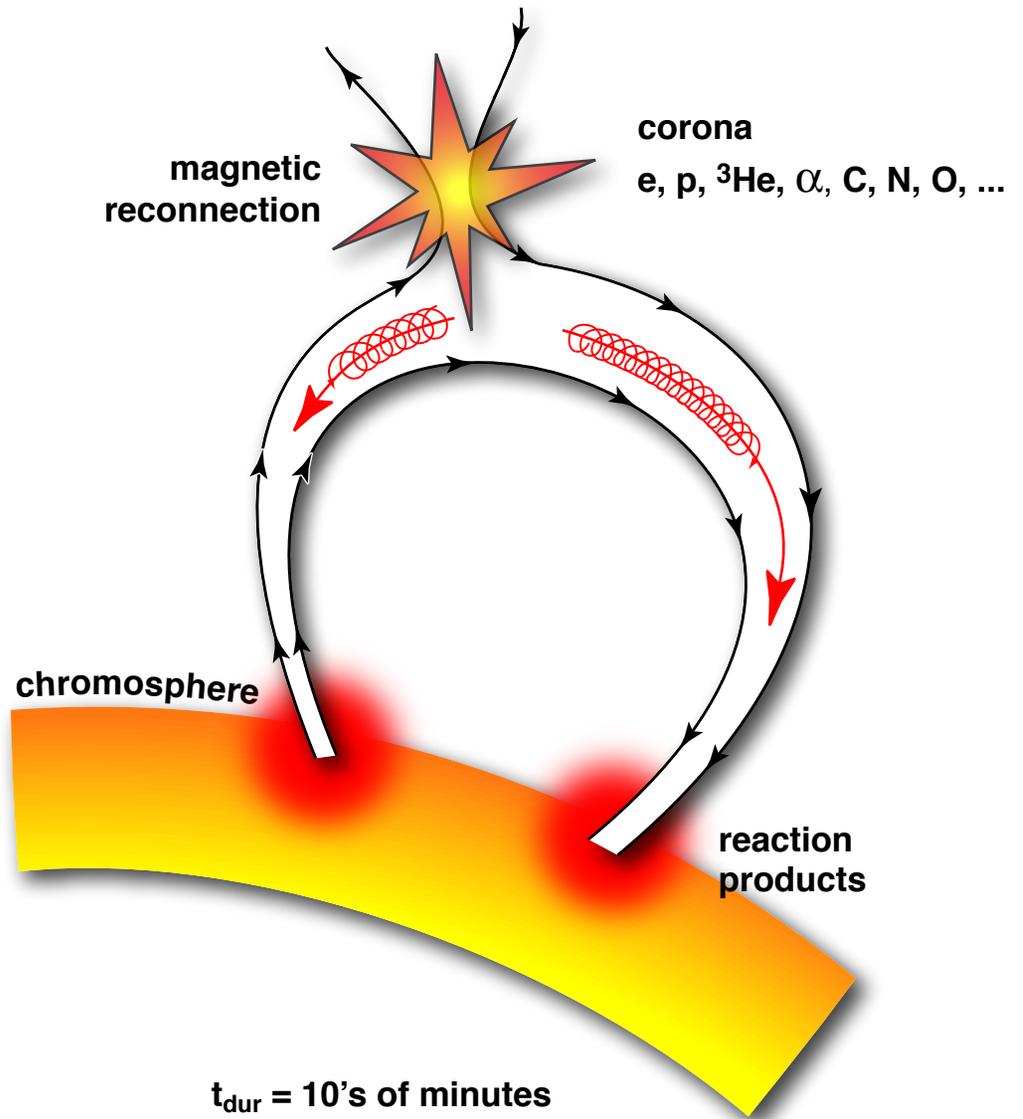
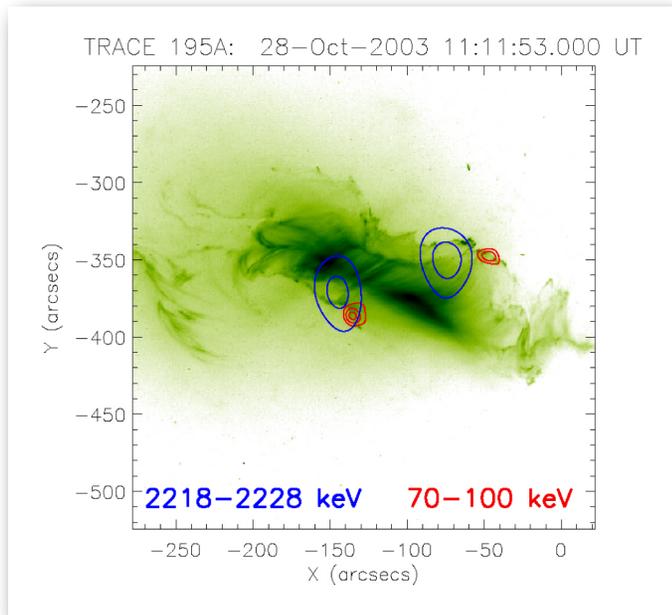


γ -ray and Neutron Production in Solar Flares

Yohkoh



RHESSI



Products of Accelerated-Particle Interactions

Products of ion nuclear interactions

excited nuclei

radioactive nuclei

neutrons

π^+ , π^- , π^0

Observable emission

e^- : X- and γ -ray bremsstrahlung

ions:

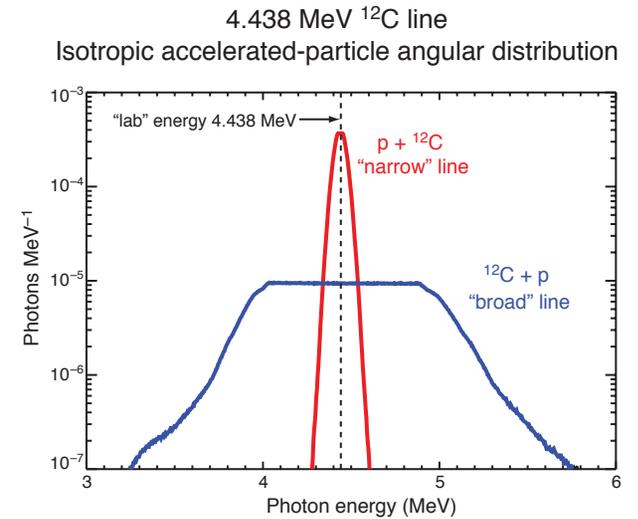
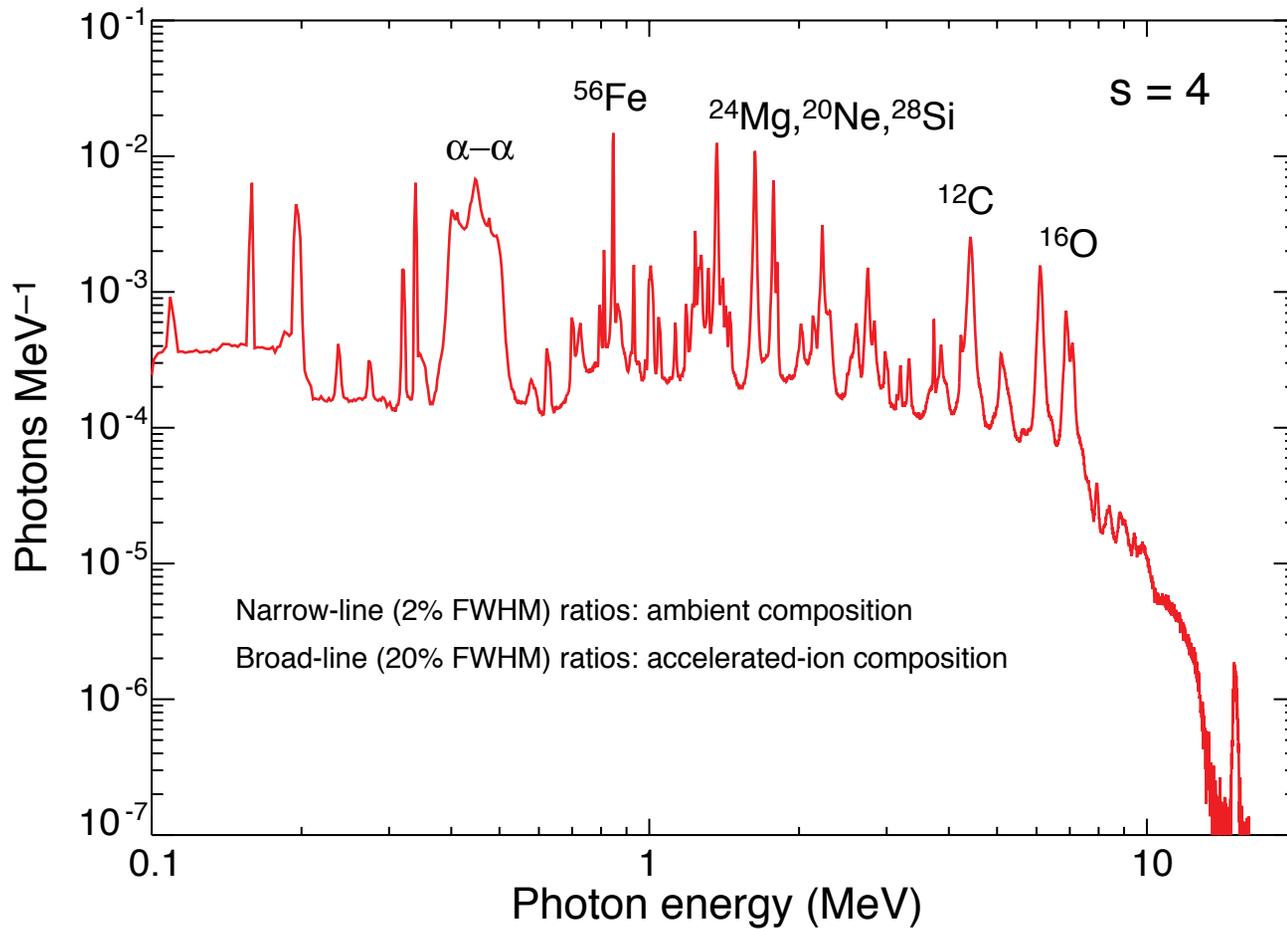
excited nuclei \rightarrow prompt γ -ray line radiation

radioactive nuclei \rightarrow $\left\{ \begin{array}{l} \text{delayed } \gamma\text{-ray line radiation} \\ e^+ \rightarrow \gamma_{511} \text{ \& continuum} \end{array} \right.$

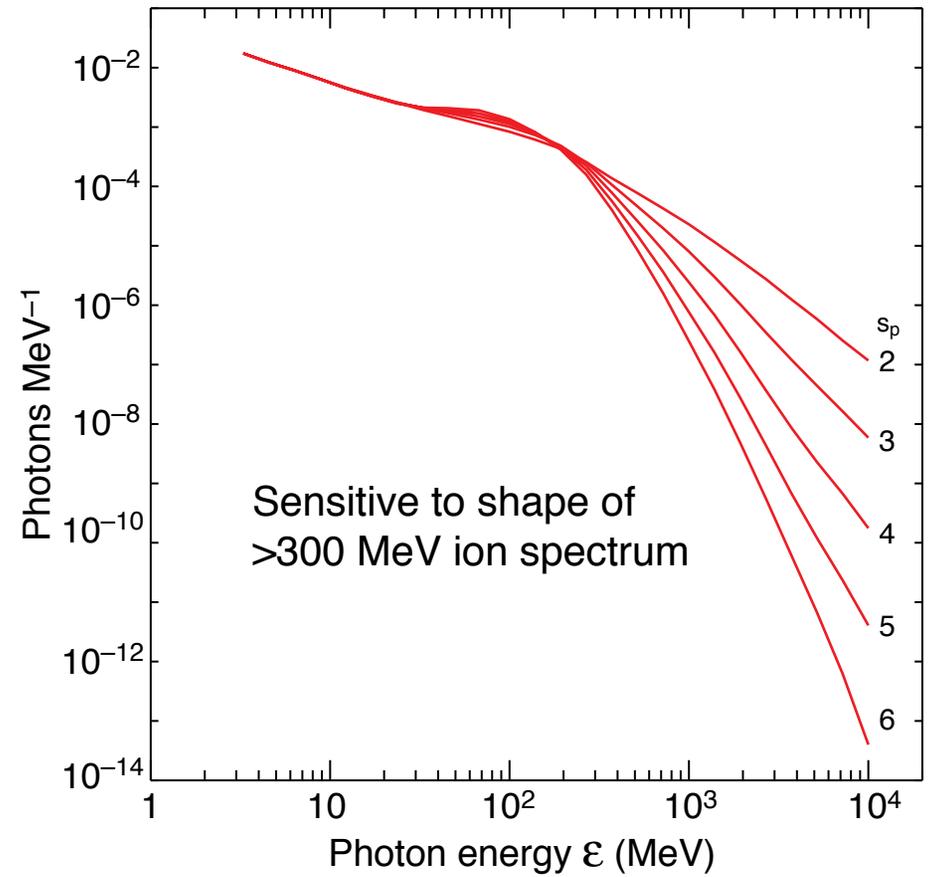
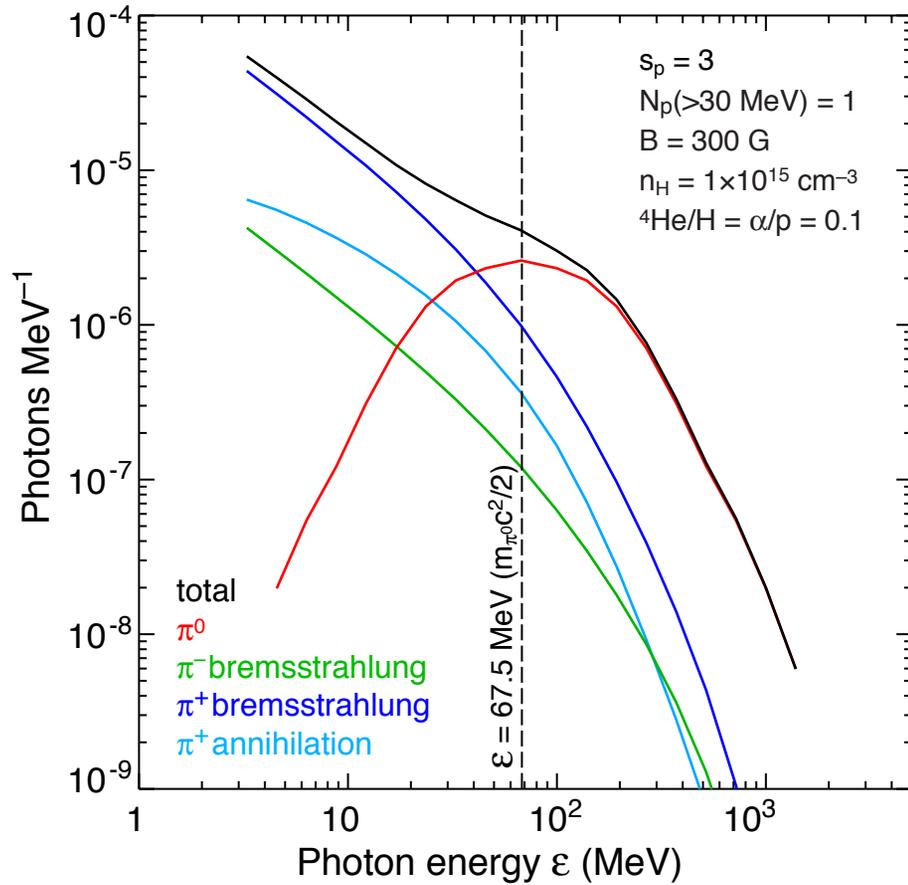
neutrons \rightarrow $\left\{ \begin{array}{l} \text{escape into space (and n-decay protons \& electrons)} \\ \text{capture on H} \rightarrow \text{d} + \gamma_{2.223} \end{array} \right.$

$\pi \rightarrow$ $\pi^0 \rightarrow 2\gamma$ $m_{\pi^0} = 135 \text{ MeV} \rightarrow \epsilon_\gamma = 67.5 \text{ MeV}$
 $\pi^{+,-} \rightarrow e^{+,-} \rightarrow \gamma_{511}$ (annihilation, line & in-flight continuum)
continuum emission via bremsstrahlung

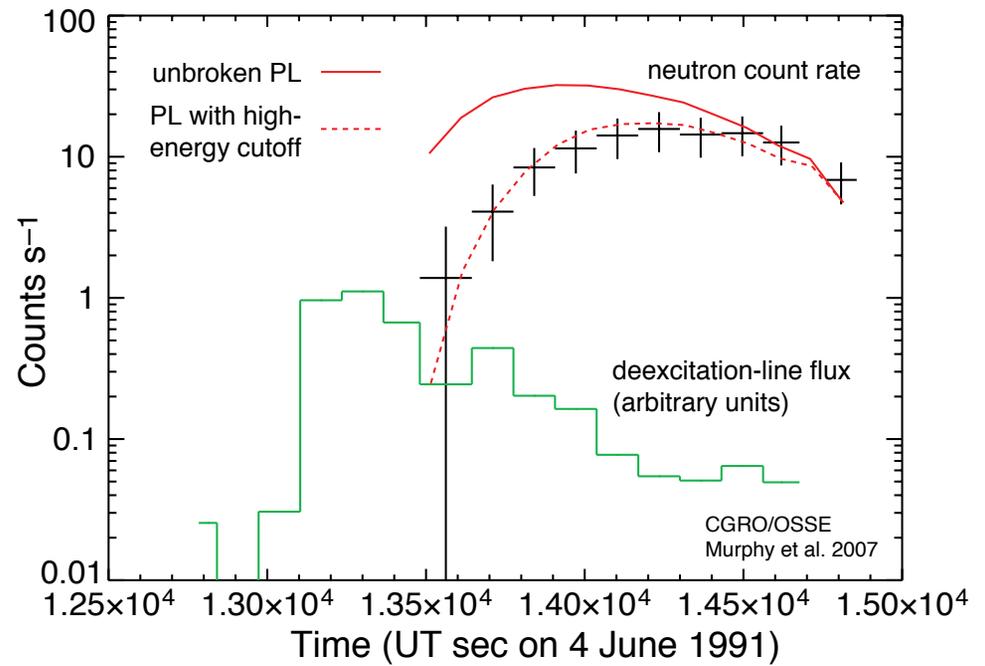
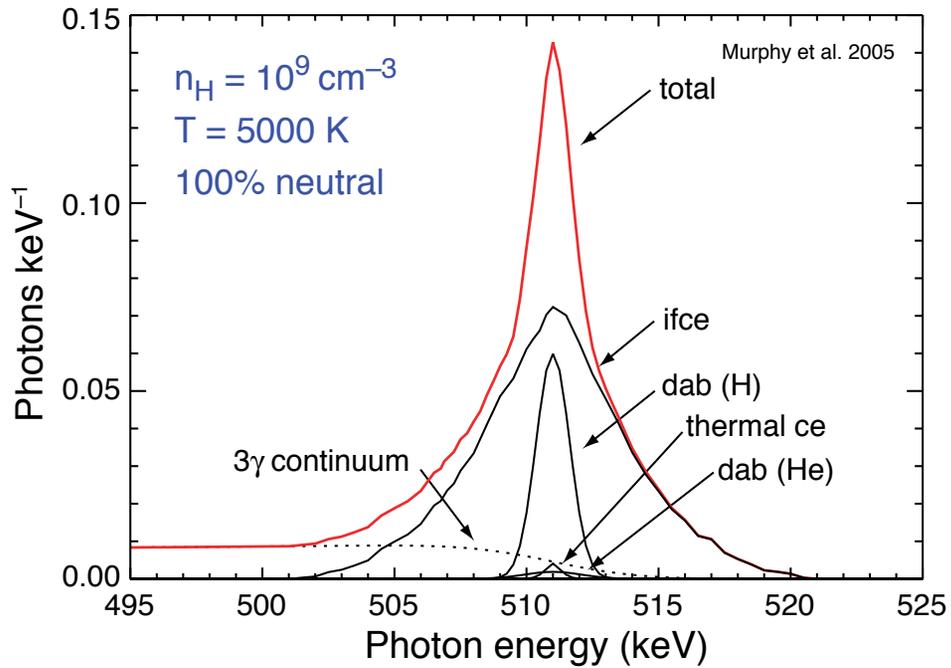
Deexcitation-line Spectrum



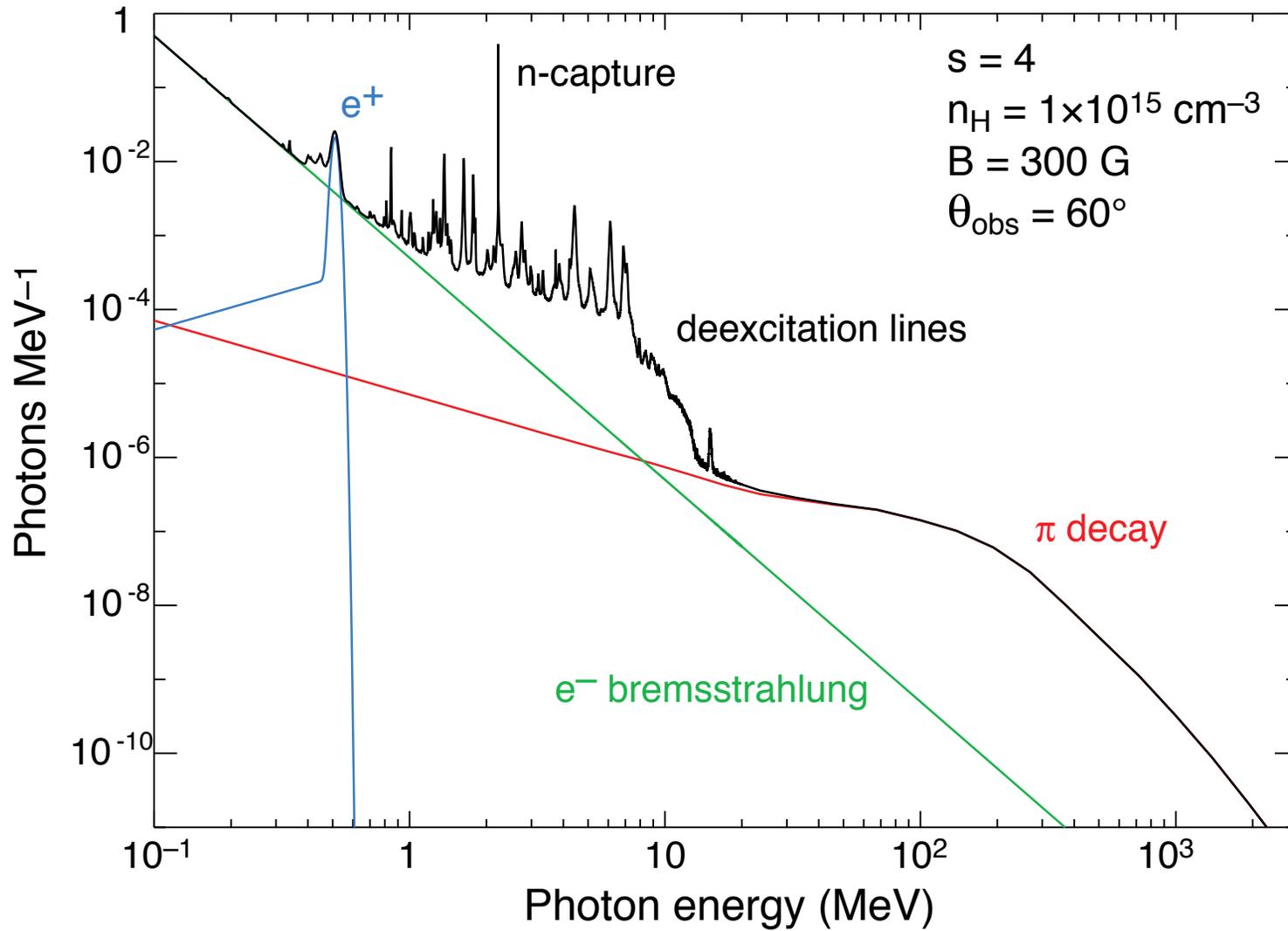
Pion-decay Spectra



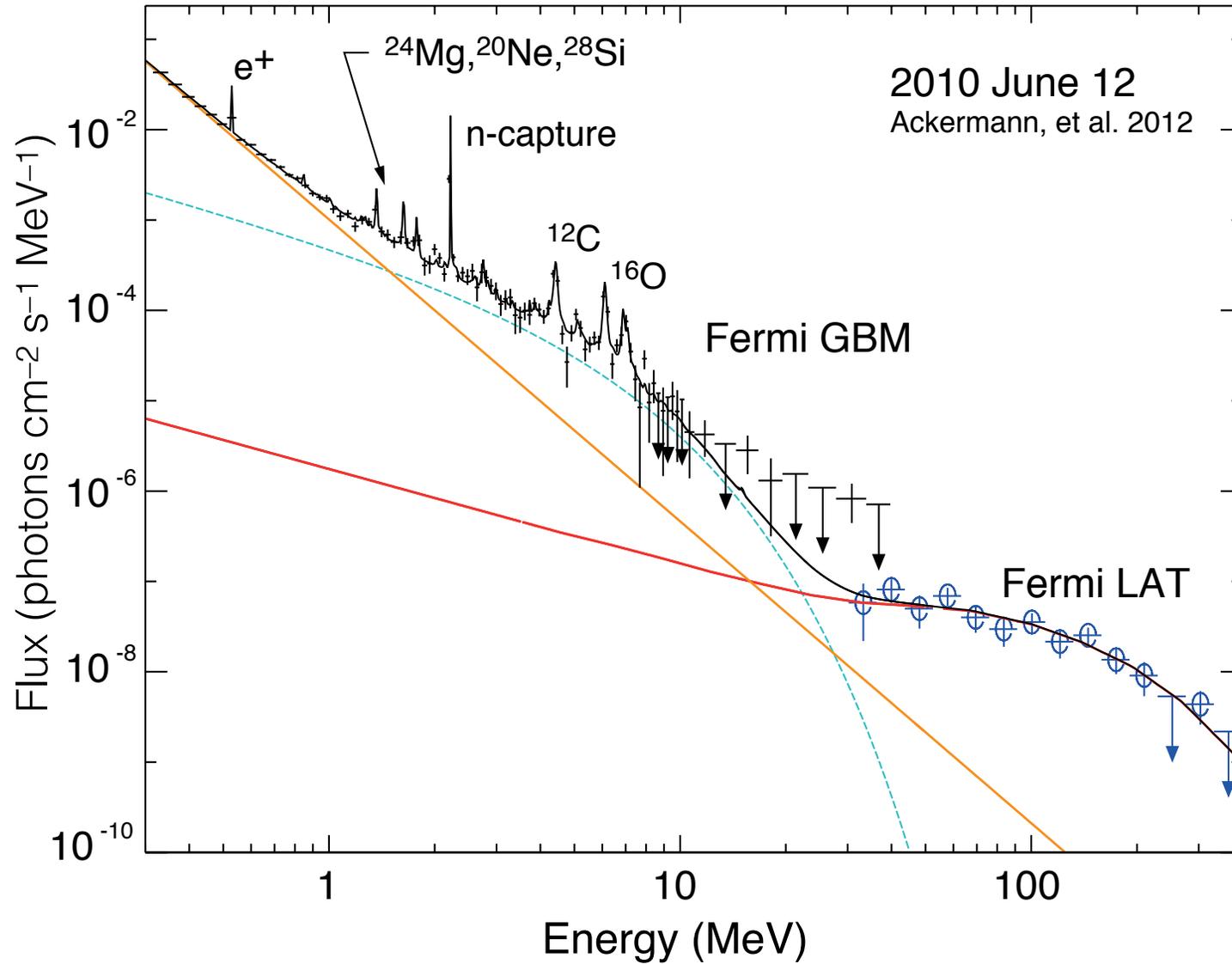
Annihilation Line and Neutrons



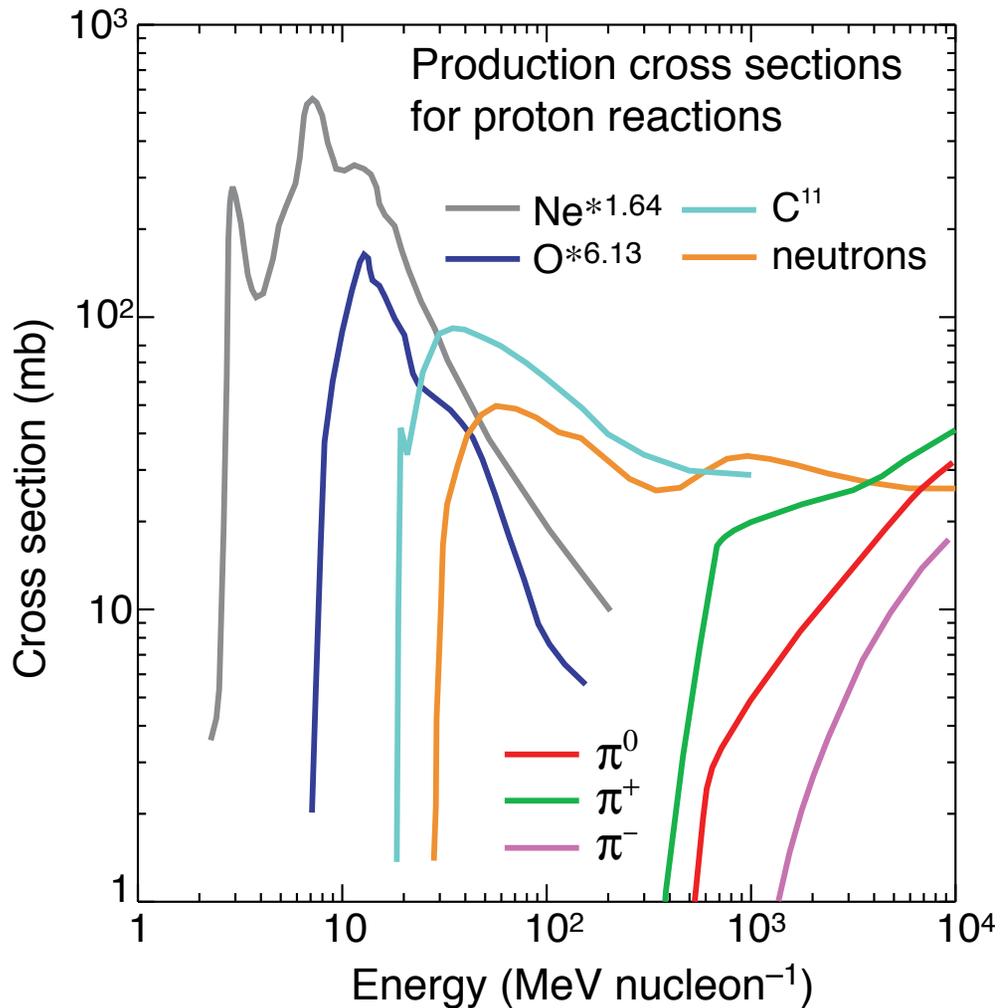
Total Gamma-ray Spectrum



Solar-flare Gamma-ray Observations



Cross Sections and Ion-energy Dependences



Radiation

deexcitation lines

n-capture

neutrons @ Earth

π -decay

Ion energy

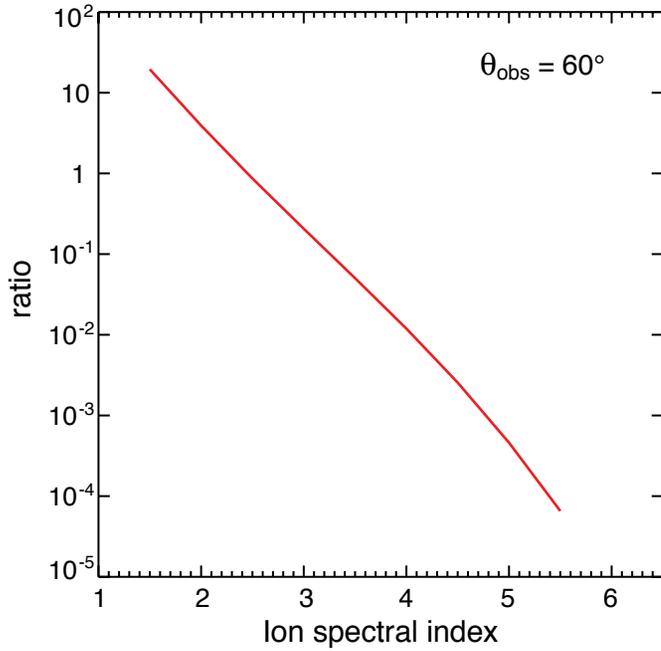
2–20 MeV nucleon⁻¹

10–30 MeV nucleon⁻¹

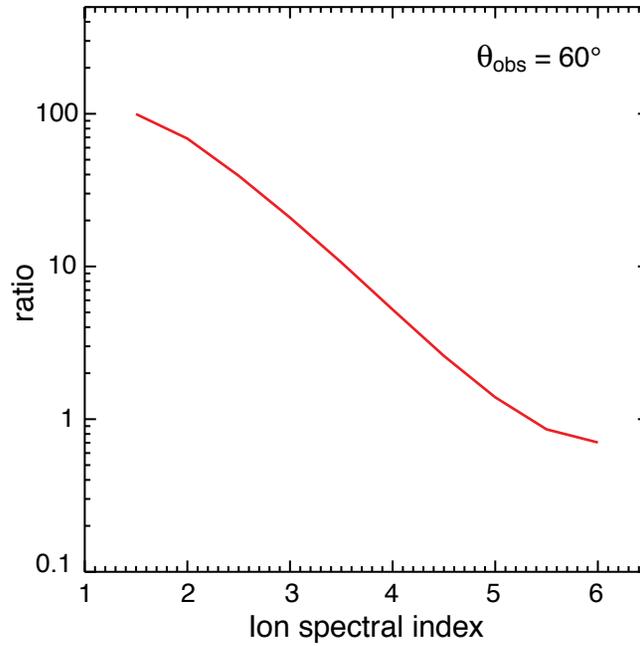
>30 MeV nucleon⁻¹

\gtrsim 300 MeV nucleon⁻¹

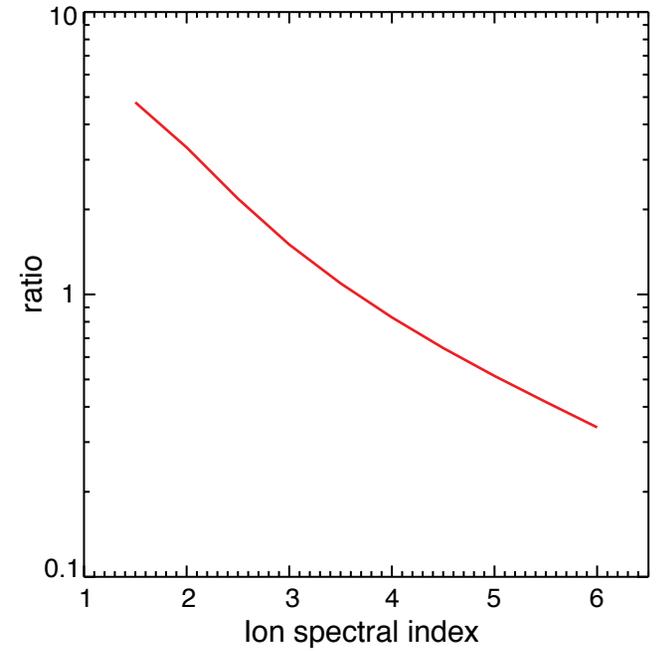
Ion Spectral Shape from Emission Ratios



$Q_{\pi(>100 \text{ MeV})}/Q_{2.2}$
[>300 MeV] — [10–30 MeV]



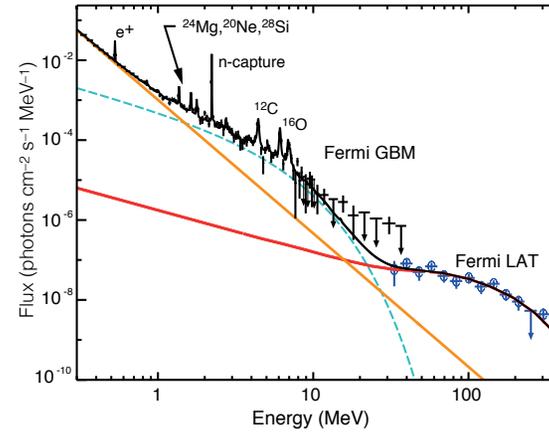
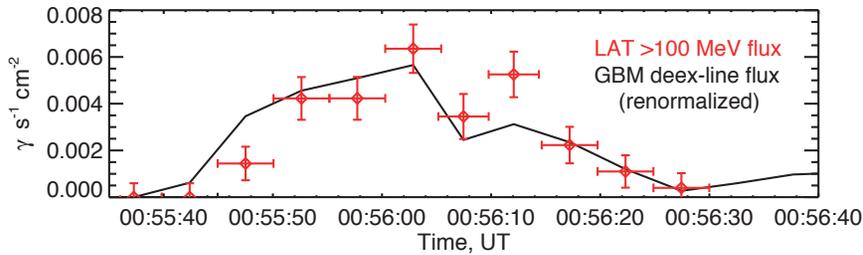
$Q_{2.2}/Q_{4.4}$
[10–30 MeV] — [5–15 MeV]



$Q_{6.1}/Q_{1.6}$
[10–20 MeV] — [2–10 MeV]

Deriving Ion Spectral Index from Fermi LAT and GBM Data

2010 June 12 flare
Ackermann, et al. 2012



$\frac{\text{n-capture}}{\text{deex lines}}$

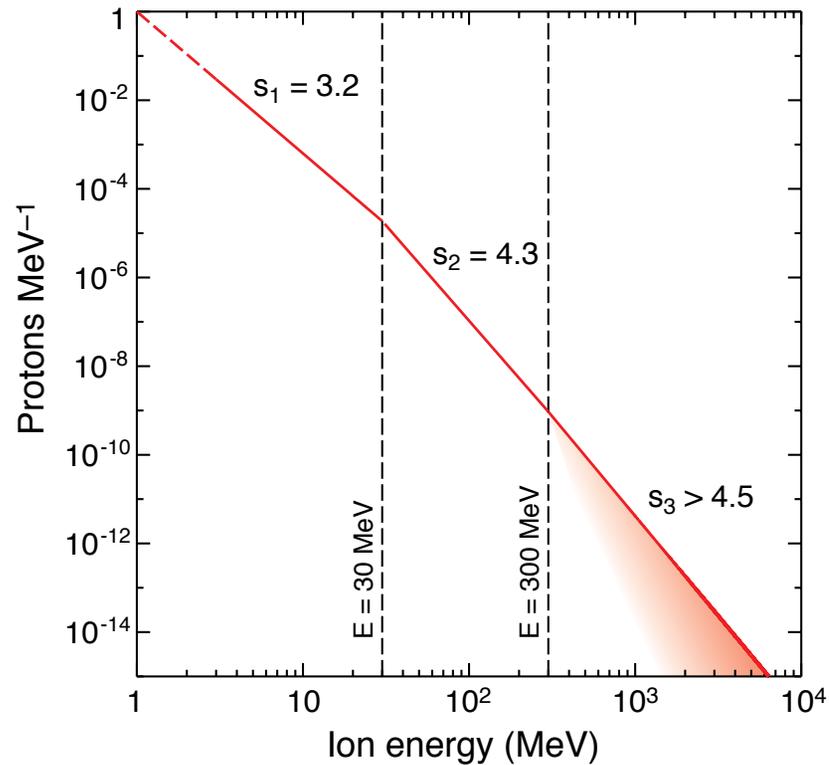
$$s_1 = 3.2$$

$\frac{\pi\text{-decay}}{\text{n-capture}}$

$$s_2 = 4.3$$

$\pi\text{-decay}$
spectral
shape

$$s_3 > 4.5$$



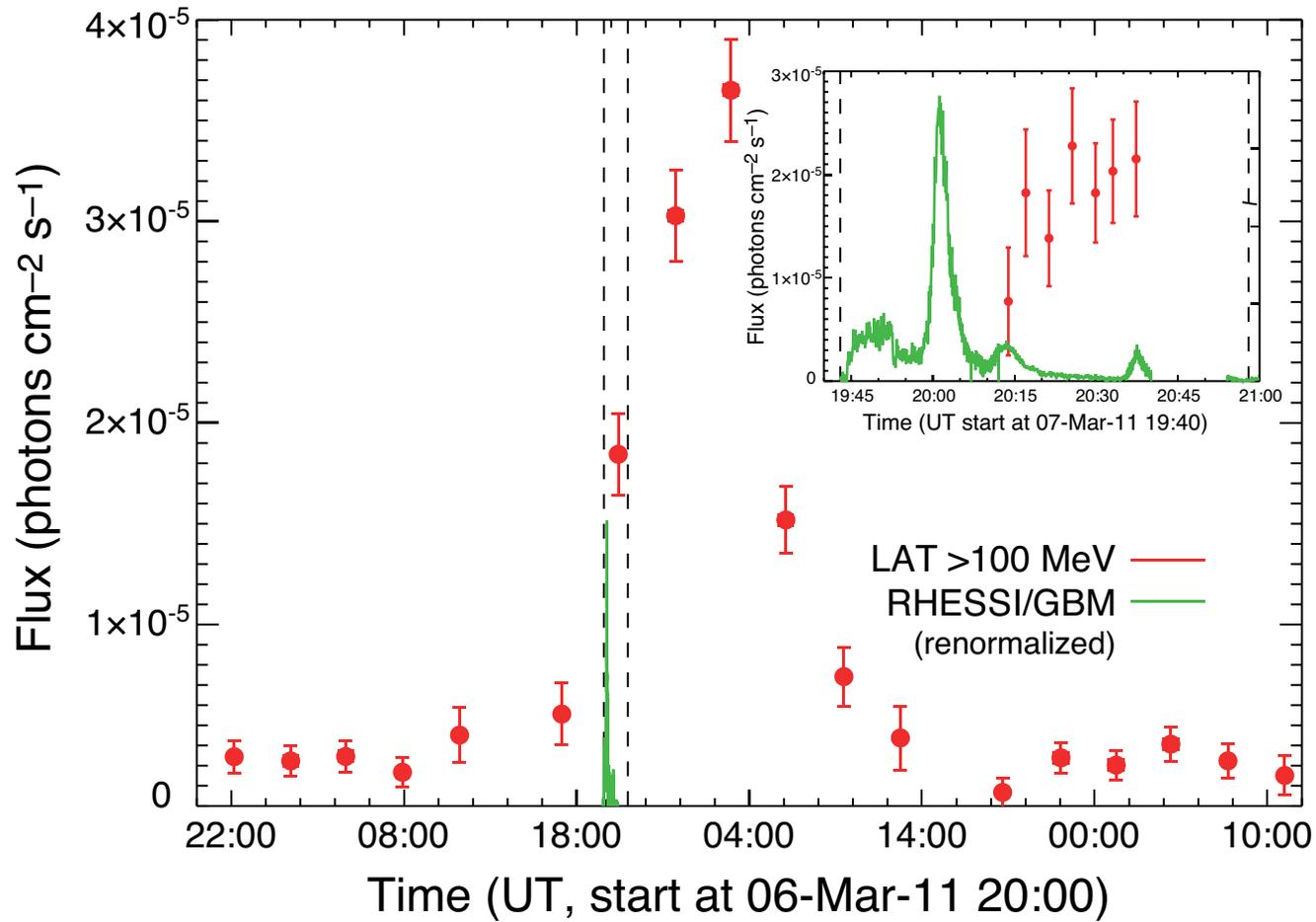
Long-duration Events

High-energy emission continues long after the impulsive phase

2011 March 7 flare

Ackermann, et al. 2014

G. Share (priv. comm.)



Source of Long-duration Emission

- continuous acceleration in a flare loop
- impulsive acceleration – trapping
- CME/shock acceleration – precipitation

2012 Mar 7 emission for 20 hr.
Ajello, et al. 2014

2013 Oct 11 over-the-limb flare
Pesce-Rollins, et al. 2015 (in prep.)

